A LOOK AT PROPAGATION FOR THE 2017/2018 CONTEST SEASON

Frank Donovan W3LPL

Carl Luetzelschwab K9LA

THANKS TO THE WWROF FOR SPONSORING THIS WEBINAR

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And thanks to Ken K4ZW for facilitating this Webinar

AGENDA

- Cycle 24, Cycle 25, solar minimums
- Disturbances to propagation
- Mitigation for disturbances
- Band-by-band summary
- DX contest strategies as Cycle 24 declines
- K4ZW's XW dilemma

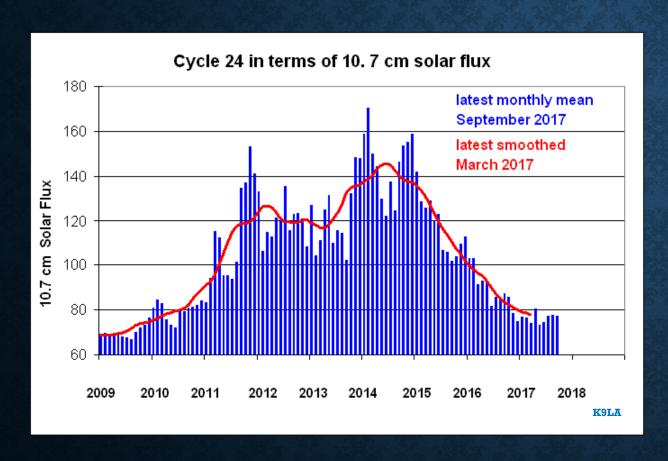
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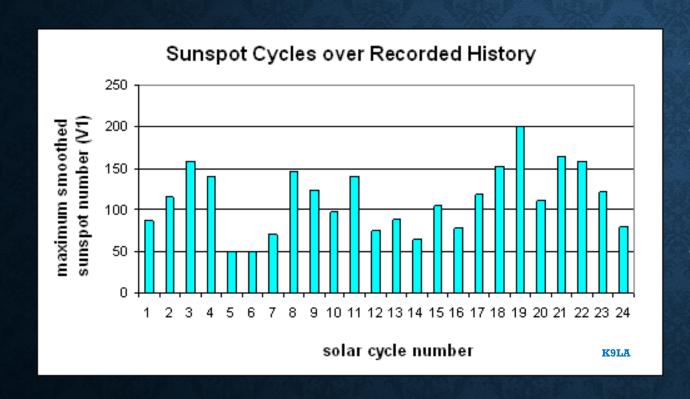
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CYCLE 24 STATUS



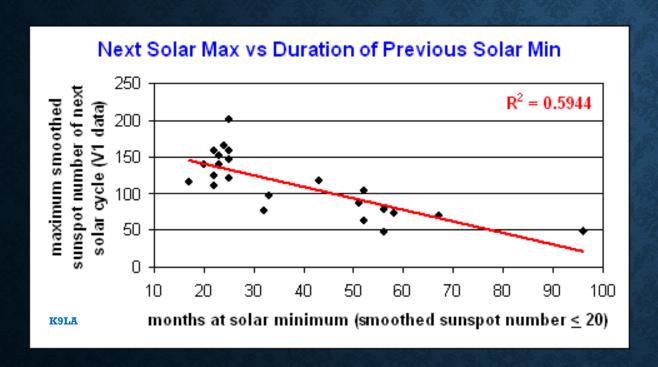
- Smallest in our lifetimes
- We still enjoyed the short second peak of Cycle 24
 - Even had 6m F2 openings
- Still have about 2 years to solar minimum

OUR PREDICTION FOR CYCLE 25



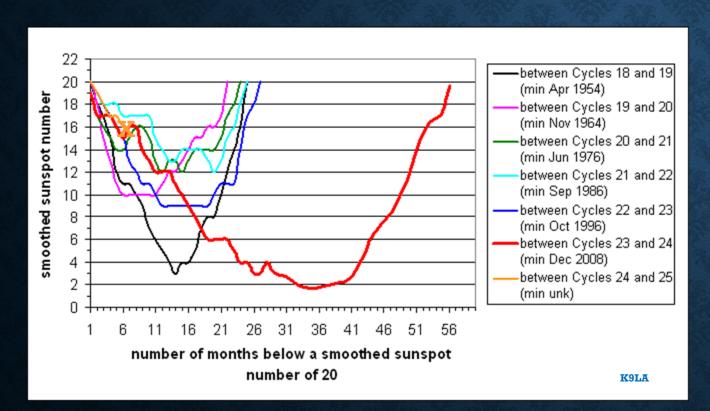
- 3 periods of big solar cycles
- 2 periods of small solar cycles
- It appears that we're entering another period of small cycles
- Solar scientists so far are in agreement

DURATION OF SOLAR MINIMUM



- The longer we're at solar minimum, the smaller the next cycle
- Conversely, the smaller the cycle, the longer the solar minimum preceding that cycle

WE'VE BEEN SPOILED



- Most of our solar minimums were about 2 years . . .
- Except for the minimum prior to Cycle 24
 - Almost 5 years!
- We may be in for another long solar minimum since
 Cycle 25 is expected to be small

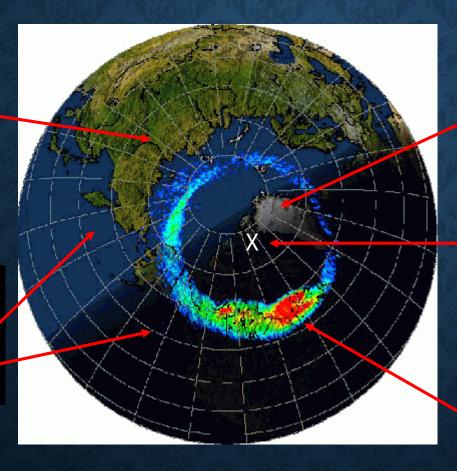
DISTURBANCES - THE BIG PICTURE

R = Radio blackout

X-ray radiation from M or X class solar flare can result in more absorption on the daylight side of the Earth

G = Geomagnetic storm

Elevated K indices can decrease F region MUFs at high and mid latitudes both day and night – but can enhance low latitude ionization



S = Solar radiation storm

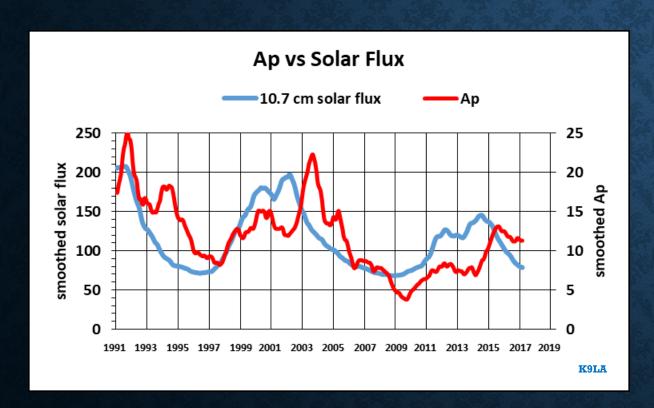
Protons from M or X class solar flare can result in more absorption in the polar cap

X marks the spot – the North magnetic pole

G = Geomagnetic storm

Elevated K indices can cause increased auroral ionization (more absorption and VHF aurora)

GEOMAGNETIC ACTIVITY VS SOLAR CYCLE



- Ap highest (most disturbed)
 at solar max and especially
 during the declining phase
- Ap lowest (quietest) at solar min and initial ascent
- Frank will have more on this

MITIGATION TO DISTURBANCES

- Geomagnetic storms (most of the time a depleted F2 region)
 - Cause: CMEs (most prevalent at solar max) and coronal holes (declining phase)
 - Mitigation: QSY to lower freqs, look for enhanced low latitude paths, look for skewed paths on the lower bands
- Solar radiation storms (energetic protons into the polar cap)
 - Cause: M and X-Class solar flares (most prevalent at solar max)
 - Mitigation: avoid polar paths, look other way around (long vs short path)
- Radio blackouts (radiation at X-ray wavelengths)
 - Cause: M and X-Class solar flares (most prevalent at solar max)
 - Mitigation: QSY to higher freqs (less loss & possible enhancement), look for paths in darkness

W3LPL

ADAPTING YOUR DX CONTEST STRATEGIES TO THE STEADILY DECLINING SOLAR CYCLE

- Spotless days are now much more frequent
 - Probably more than 100 spotless days this year
 - Many more spotless days during the next five years
- Solar flux will continue its decline to much lower levels
 - 60s and 70s will be increasingly common for at least five years
- Frequent geomagnetic disturbances until Solar Minimum
 - at least until the onset of solar minimum in about 2019
 - Mostly driven by frequent -- often intense -- coronal hole activity
 - Solar flares and coronal mass ejections (CMEs) will be much less frequent
 - Geomagnetic disturbances will be much less frequent after solar minimum in about 2019

MAJOR IMPROVEMENT IN 160 METER WORLDWIDE PROPAGATION

- Stronger DX signals
- More reliable DX openings especially to Europe and Japan
- Lower absorption
 - During increasingly common periods of less intense geomagnetic activity
 - Especially for a few years after solar minimum during about 2020-2022
- Less daytime D layer absorption
 - DX openings begin just before sunset and continue just after sunrise
- Less intense night time E layer absorption as the E layer MUF drops

MAJOR IMPROVEMENT IN 160 METER WORLDWIDE PROPAGATION

Continuous openings to Europe, Mid-east and north Africa 2200-0830Z

• Frequent strong JA openings at sunrise mid-Nov to mid-Feb 1200-1230Z

• More frequent direct short path polar opening to central Asia 2300-0100Z

- More crowded band conditions
 - Especially when there is no strong 40 and 80 meter propagation to Europe

SIGNIFICANTLY IMPROVED 80 METER WORLDWIDE PROPAGATION

- Stronger DX signals
- Much more reliable DX openings especially to Europe and Japan
- Consistently low absorption
 - During increasingly common periods of less intense geomagnetic activity
 - Especially for a few years after solar minimum during about 2020-2022
- Less daytime D layer absorption
 - DX openings begin well before sunset and continue well after sunrise

SIGNIFICANTLY IMPROVED 80 METER WORLDWIDE PROPAGATION

- Continuous east coast USA openings to Europe, Mid-east and north Africa
 - 2130-0830Z
- Regular strong east coast USA to JA openings at sunrise mid-Nov to late Feb
 - 1130-1300Z
- Direct short path polar opening to central Asia will be much more frequent
 - 2130-0200Z
- Much more crowded band conditions
 - Especially when there is no strong 40 meter propagation to Europe

NEARLY 24 HOUR WORLDWIDE 40 METER PROPAGATION

- Much less daytime D layer absorption
- Nearly 24 hour DX openings during CQWW CW in late November
- DX openings to Europe will often begin before east coast USA mid-afternoon
 - Especially from mid-November through late February
- But DX openings to Europe will often end shortly after sunset
 - Especially from mid-November through late February
 - Direct short path to Europe often opens again at European sunrise
- Trans-polar propagation will become much less frequent
 - East coast US to East Asia propagation will often be via the South West skew path
 - Strong direct short path often opens at east coast sunrise

NEARLY 24 HOUR WORLDWIDE 40 METER PROPAGATION

•	Europe, Mid-East an	d north Afri	ica prop	agation

•	Activity QSYs to 40 meters before mid-afternoon	~1930Z
	Don't miss the strong mid oftenness /exemines an exinera	2000 020

•	Propagation often fades.	fails a few ho	urs after sunset	0200-0600Z
			SHOULD BE STONE TO AN ADDRESS OF THE SHOULD	

- Strong openings usually resume at sunrise in Europe ~0600-0900Z
- Japan, Far East and Central Asia propagation

•	Brief direct s	short pa	ath ope	ning a	t JA sunset	0800-0900Z
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- Weak skew path opening at about 240° azimuth ~0900-1130Z
- Strongest short path JA opening from the east coast ~1130-1300Z
 Strong long path Asia signals at 150° azimuth 2130-2215Z
- VK/ZL and southeast Asia long path at 90-150° azimuth 2100-2300Z
- Southeast and central Asia long path at ~240° azimuth ~1130-1300Z

SHORTER DURATION WORLDWIDE 20 METER PROPAGATION

- Mostly a daytime DX band from November to late February
- All DX propagation usually ends well before midnight on the east coast
 - Especially from mid-November through late February
- DX openings to Europe usually begin about an hour before east coast sunrise
- But DX openings to Europe usually end shortly after sunset in Europe
 - Especially from mid-November through late February
- Much shorter duration east coast USA propagation to east and central Asia
- US to central and east Asia propagation will often be via the South West skew path
 - Especially from 1200-1400Z at 240 degrees azimuth

SHORTER DURATION WORLDWIDE 20 METER PROPAGATION

All DX propagation usually ends well before midnight ~0300Z
Sporadic, weak night time Africa & south Pacific openings 0500-0700Z

• Europe, Mid-East and north Africa propagation

• from before east coast USA sunrise until mid-afternoon ~1000-1900Z

• the opening is sometimes delayed until after sunrise

• the opening ends much earlier in the afternoon than in recent years

• Japan, Far East and central Asia propagation

• short evening short path opening 2100-0100Z

• morning short path opening 1300-1500Z

• both openings are much shorter than in recent years

• South Asia and Mid-East morning long path 1300-1500Z

• VK, ZL and south Pacific mid-afternoon long path 1900-2200Z

MUCH SHORTER DURATION WORLDWIDE 15 METER PROPAGATION

- <u>Usually</u> a daytime only DX band from November to late February
- All DX propagation usually ends shortly after sunset on the east coast
 - Especially from mid-November through late February
- DX openings to Europe usually begin about an hour after east coast sunrise
- But DX openings to Europe usually end shortly before sunset in Europe
 - Especially from mid-November through late February
- Very short duration east coast USA propagation to east and central Asia
 - Sometimes via the long path at 150 degrees azimuth

MUCH SHORTER DURATION WORLDWIDE 15 METER PROPAGATION

- Europe, Mid-East and north Africa propagation
 - from about an hour after our sunrise until early afternoon 1300-1800Z
 - shorter openings than we've enjoyed in recent years
- Japan and Far East propagation
 - weak unreliable late afternoon short path opening 2130-2300Z
 - sometimes only via long path to 150 degrees azimuth 1300-1400Z
 - much shorter openings than we've enjoyed in recent years
- 15M usually closes shortly after our sunset ~2200Z
 - <u>always</u> stays closed all night

NO LONGER A WORLDWIDE DX BAND 10 METER PROPAGATION

- Strictly a daytime only DX band
- All DX propagation ends before sunset
- There are no short path DX openings to Europe
 - There are weak scatter openings at 140 degrees azimuth
- But DX openings to Europe usually end shortly before sunset in Europe
 - Especially from mid-November through late February
- There are no short path DX openings to Asia
 - Rare very weak signal scatter openings to east Asia at 230 degrees azimuth

NO LONGER A WORLDWIDE DX BAND 10 METER PROPAGATION

- South America, Caribbean and Central America
 - PY and LU activity has increased significantly in recent years
 - usually opens about an hour after our sunrise

~1300Z

- opening can fade for an hour or two, then return much stronger in the afternoon
- usually closes about an hour before our sunset

~2100Z

- Europe and north Africa
 - very weak signals on the SE skew path at 110-150° azimuth ~1400-1700Z
- VK/ZL and South Pacific
 - A fairly reliable weak signal opening

~1900-2100Z

- Japan, North Pacific and Far East
 - rare morning weak signal long path at 150° azimuth

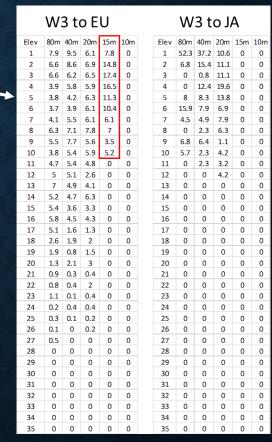
~1300-1400Z

• rare evening very weak signal skew path 200-240° azimuth

2100-2200Z

DX CONTEST STRATEGIES FOR DECLINING LOW SOLAR ACTIVITY

- High antennas are much more important during solar minimum
- Improve your low band transmitting and receiving antennas!
- Start every DX contest on 40 meters
 - The European opening begins before mid-afternoon
 - the strong European opening often ends a few hours after sunset



DX CONTEST STRATEGIES FOR DECLINING LOW SOLAR ACTIVITY

 Capitalize on improved 160 and 80M propagation 	2200-0830Z			
 especially important when 40 meters is not strongly open 	 especially important when 40 meters is not strongly open to Europe 			
 Strong 40 meter opening after sunrise in Europe 	0600- <mark>0900</mark> Z			
• 160, 80 and 40 meter openings to VK, ZL and JA	0900-1230Z			
• 20M European opening starts before sunrise	1000-1900Z			
 15M European opening starts about an hour after sunris 	se 1200-1800Z			
 10 meters opens primarily to the south 	1300-2100Z			
 Strong 40M afternoon/evening openings to Europe 	2000-0300Z			
 20 meter evening openings to Japan 	2100-0100Z			

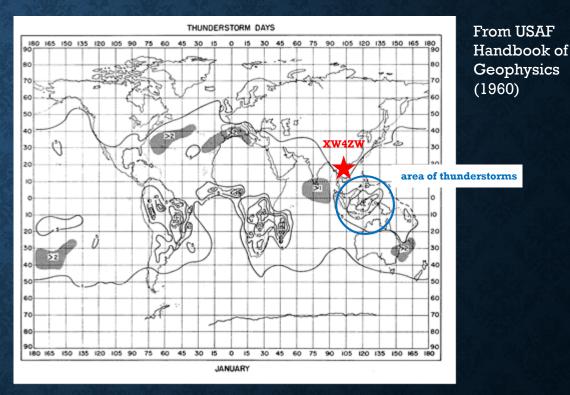
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K4ZW'S DXPEDITIONS TO XW

- XW4ZW January 23-30, 2015
 - 159 NA on 80m, 8 NA on 160m, 1947 total on 80m and 160m
- XW4ZW January 16-22, 2016
 - 49 NA on 80m, 11 NA on 160m, 987 total on 80m and 160m
- XW4ZW January 9-15, 2017
 - 52 NA on 80m, 1 NA on 160m, 1233 total on 80m and 160m
- XW to NA is a tough path

K4ZW'S OBSERVATIONS

- Ken says it's pretty quiet to the south (long path) towards NA
- NA signals on 80m are often right at the noise
- Ken says he is being heard much better than he is hearing
- What's the problem?



If we believe the ITU Radio Noise document, XW4ZW's noise on 80m is about 20 dB worse than noise at W1 – that could contribute to non-reciprocity

MORE BEER MONEY AND A DECK OF CARDS?

- When the noise is in the same direction as your target location, there's not much you can do
- Other possible "RF" causes
 - Charly HSØZCW observes that short path <u>signal strengths</u> from HS on 20m are stronger in EU and NA than the other way
 - The author of a NATO report on the ionosphere says most paths are non-reciprocal in <u>signal strength</u> due to polarization mismatch
- Conclusion: For his next trip, Ken may have a lot of extra time for vices ©

SWEEPSTAKES

- Although we talked about DX contests, propagation for Sweepstakes and other forthcoming contests will be similar
 - Higher bands not too productive
 - Lower bands best
- All of this could also apply for the next 2-5 years
 - Depends on the duration of the next solar minimum

Q & A